

# A Radio For Hidden Photon Dark Matter

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# The Dark Matter Zoo

- What do we know?
  - Astrophysical observations strongly suggest that dark matter exists and some component of it is cold (non-relativistic).
  - Over the last two decades, there has been tremendous progress in probing WIMP dark matter phase space, and WIMP direct detection experiments are now approaching the neutrino background.
  - This motivates broadening our search and running complementary probes for other well-motivated particles.
- Is dark matter heavy ( $>1$  eV) or light ( $<1$  eV)?
- Light dark matter candidates
  - Must be boson: scalar, vector, tensor

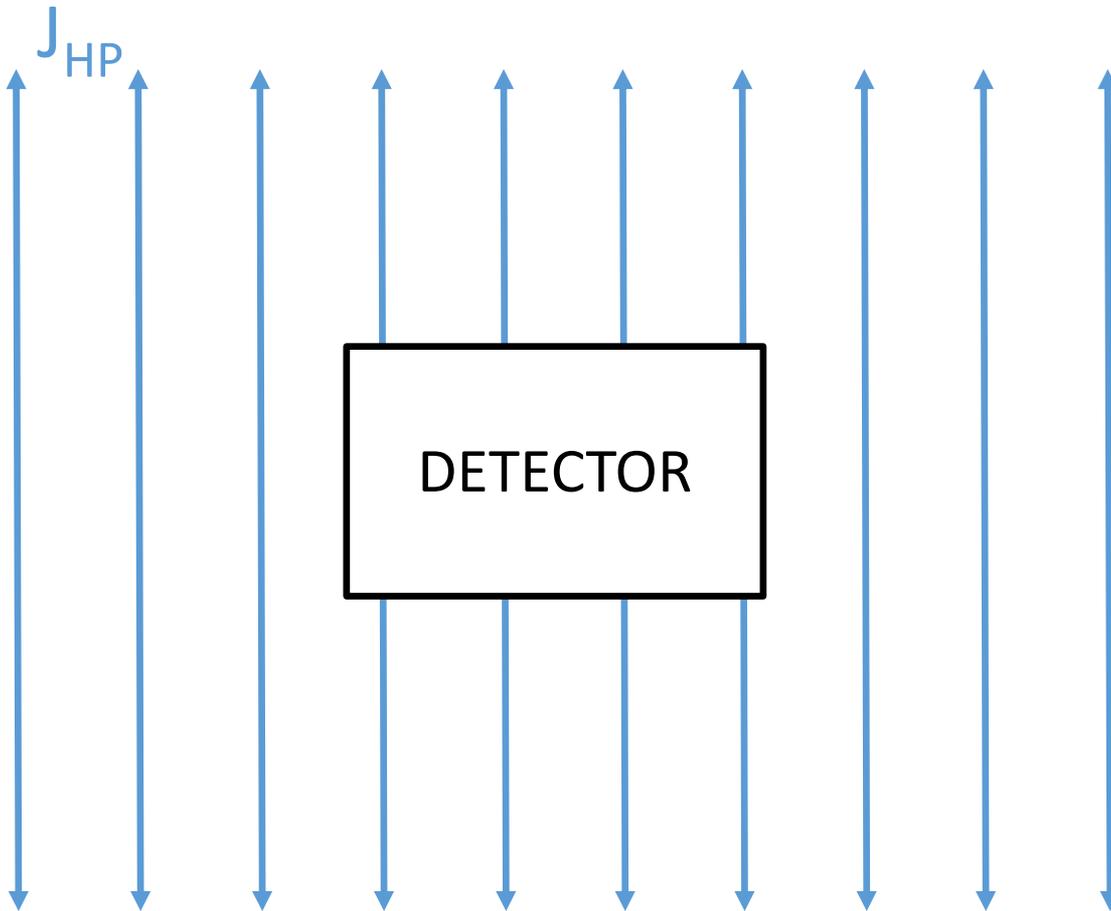
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# Hidden Photon Dark Matter: A Natural Dark Matter Candidate

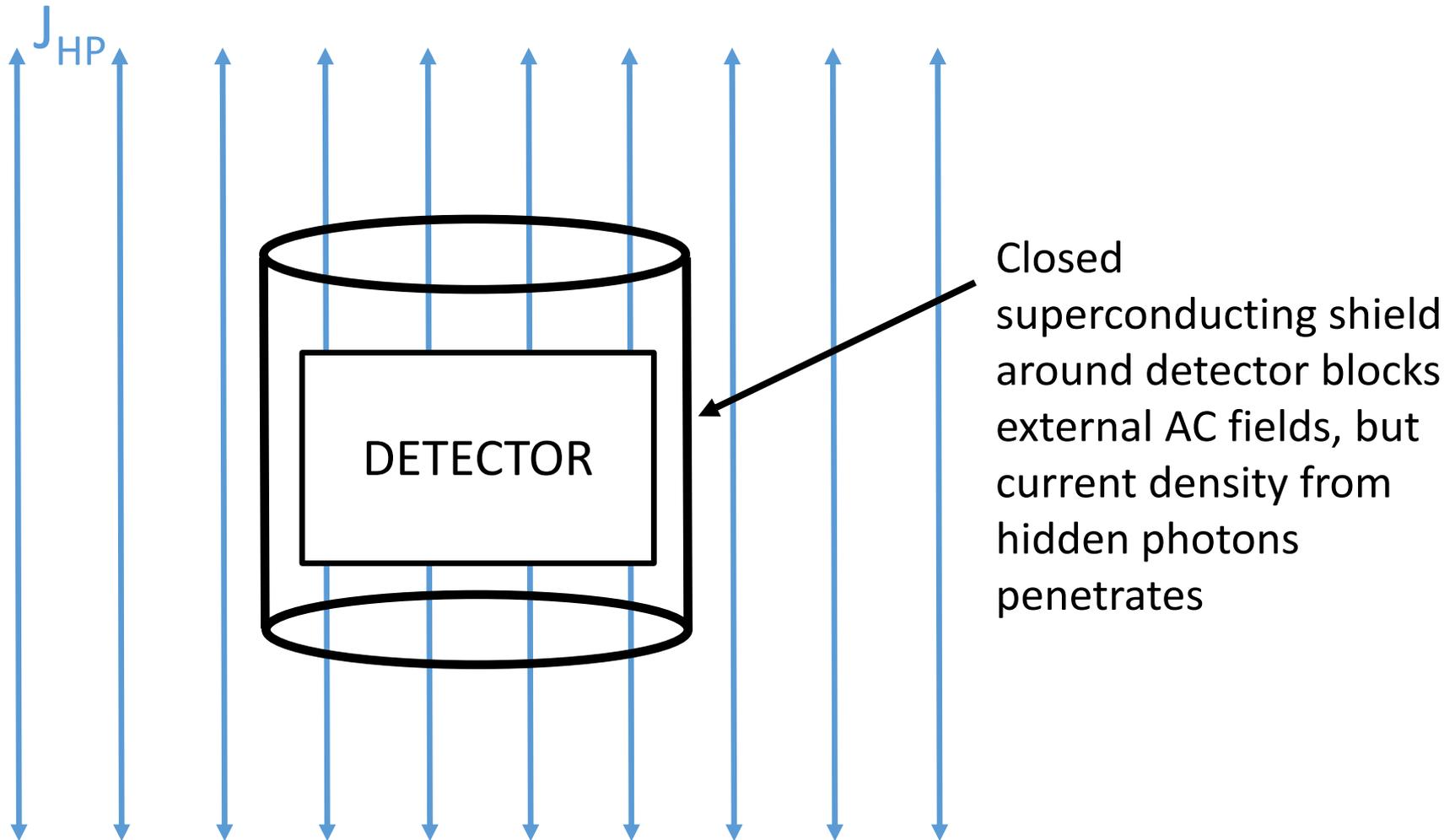
- Massive vector boson
- Couples to regular photon
- Coupling constant:  $\epsilon < 10^{-7}$
- Hidden photon is generic—naturally appears in extensions of Standard Model
- Inflation would produce hidden photons (arXiv: 1504.02102), as would other non-thermal processes
- Largely unexplored phase space – opportunity!

# How do we detect hidden photons?



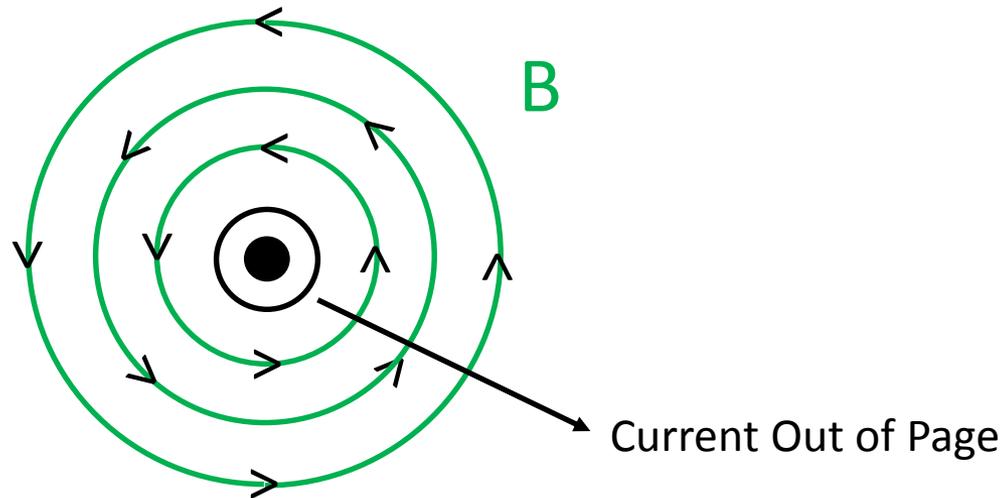
- Visualize hidden photon field as AC current density  $J_{HP}$
- Frequency of oscillation set by mass,  $hf=mc^2$  ( $\sim 100$  Hz-  $\sim 1$  THz)
- Detection in quasi-dc limit, within coherence length

# Closed Superconducting Shield Blocks External AC fields



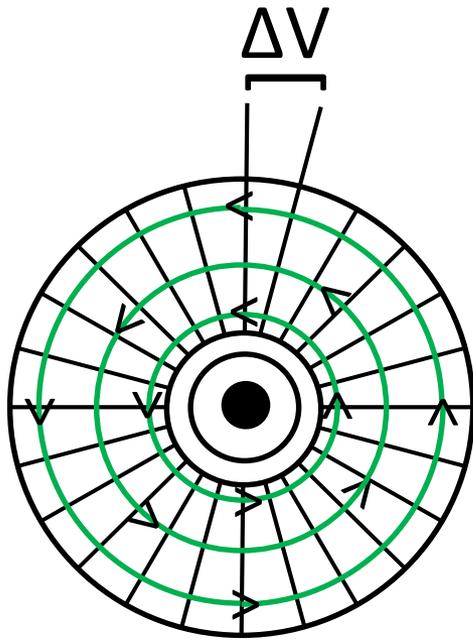
# E&M 101: How do we detect quasi-DC current without contact?

- Wire current normal to page gives circumferential quasi-DC B-field
- Question becomes: How do we pick up quasi-DC B-field?

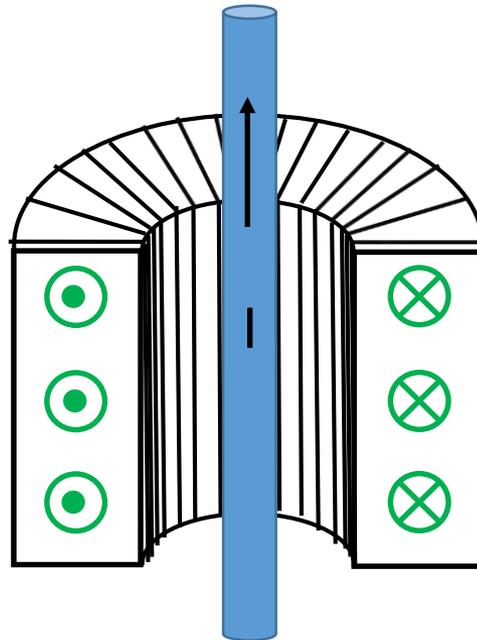


# Answer: With toroidal solenoid!

- AC magnetic field induces voltage across toroidal solenoid
- A standard idea...current clamp does the same thing!



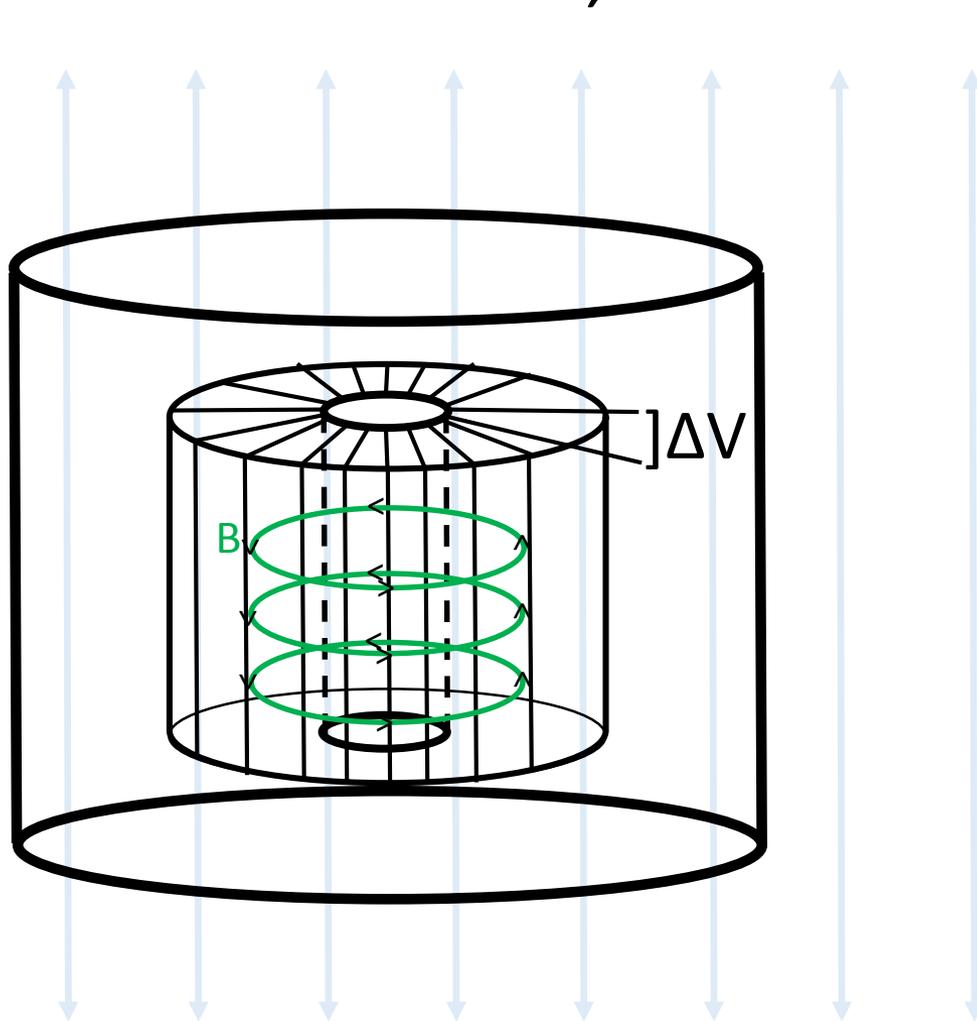
Top View



Cross-Sectional View



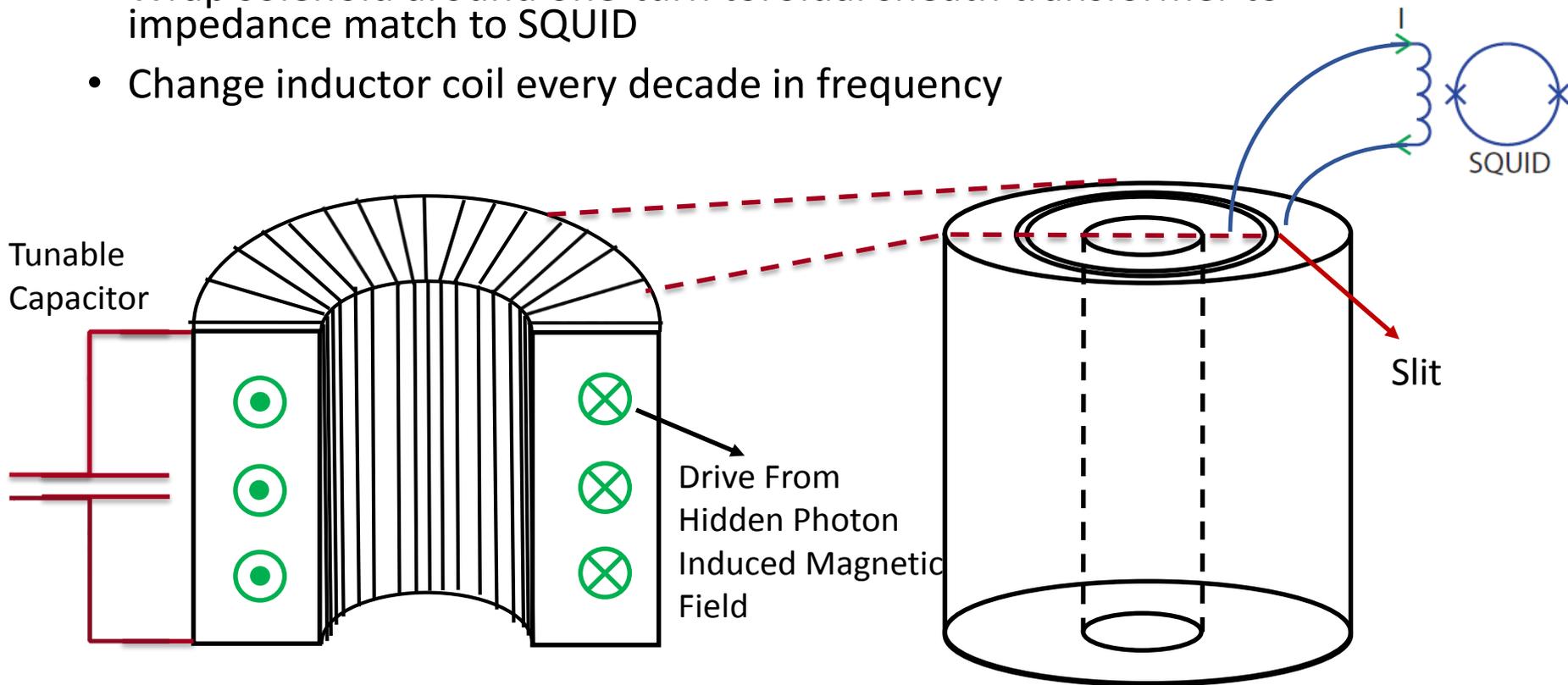
# Toroidal Solenoid Detects Hidden Photons, But Low SNR



- Current  $J_{HP}$  creates circumferential quasi-dc magnetic field inside shield
- Use toroidal geometry to pick up magnetic field
- Read out voltage with amplifier, e.g. SQUID
- Inherently broadband, but would need to integrate forever!

# Solution: Use LC resonator to get High SNR

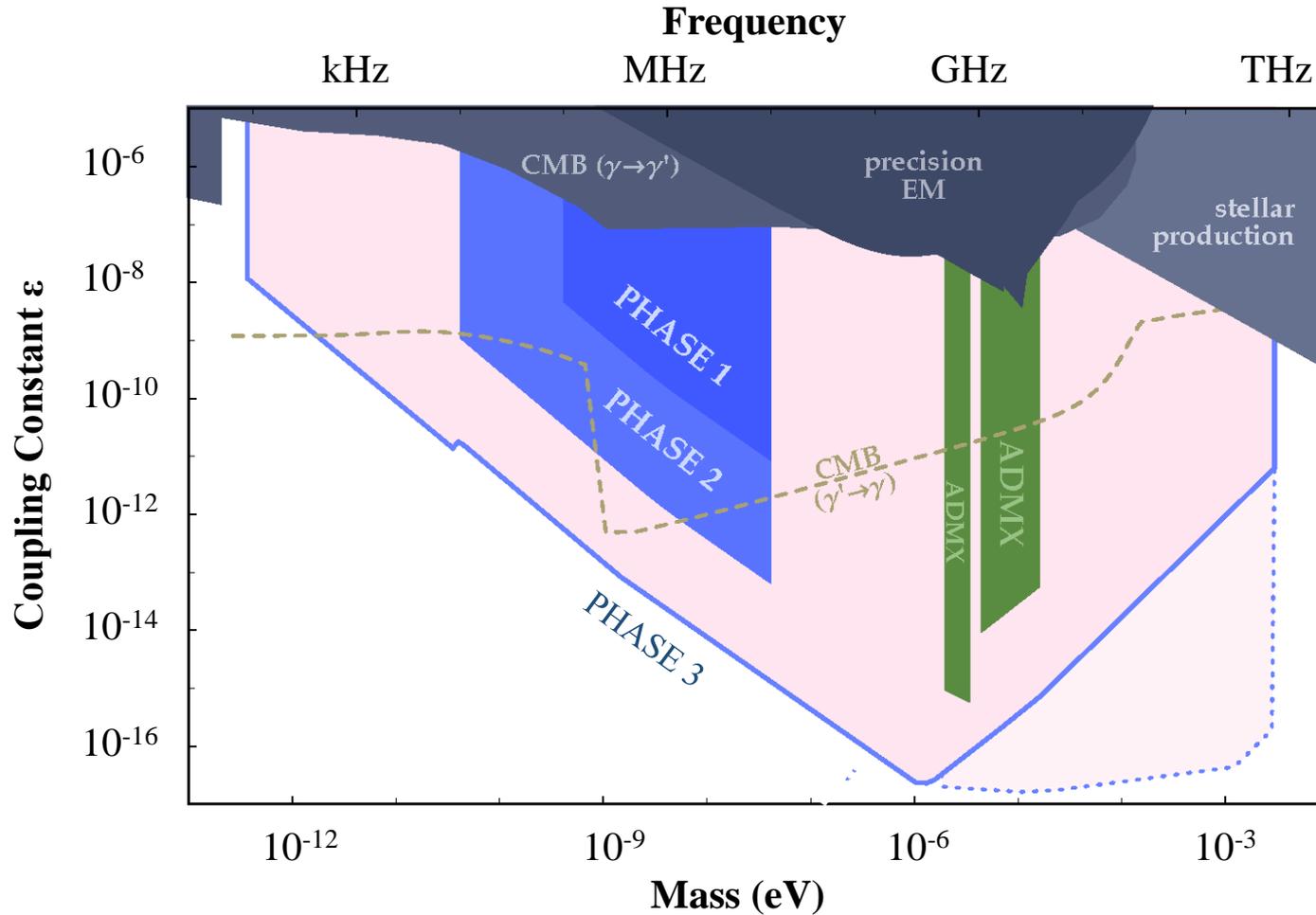
- Attach tunable capacitor to solenoid to make tunable LC resonator
- Narrowband, but high signal-to-noise
- Wrap solenoid around one-turn toroidal sheath transformer to impedance match to SQUID
- Change inductor coil every decade in frequency



# Architecture Constraints

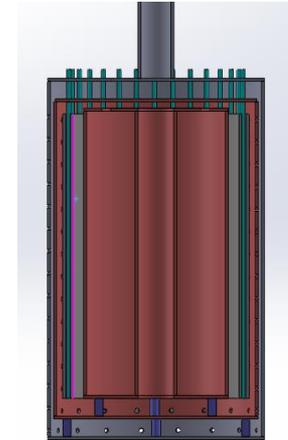
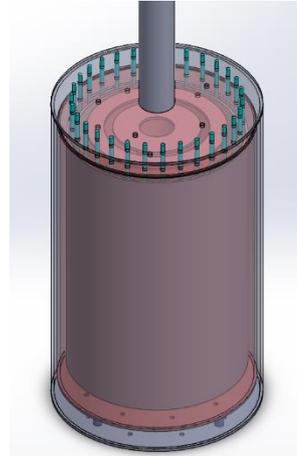
- Constraints:
  - Optimal Q of  $10^6$  set by virial velocity
  - Coupling to amplifier architecture strong enough to be resonator thermal-noise limited (low f) or quantum-noise limited (high f), but ...
  - Weak enough to make Q degradation due to amplifier backaction negligible
- Solutions:
  - Use superconducting materials (niobium, NbTi wire)
  - $f < 10$  MHz: dc SQUID
  - $10 \text{ MHz} < f < 1 \text{ GHz}$ : microwave SQUID
  - $f > 1 \text{ GHz}$ : quantum-limited parametric amplifiers (similar to ADMX), future- bolometers or single-photon counting

# Reach- 10 Orders of Magnitude in Mass, ~16 in Cross Section!

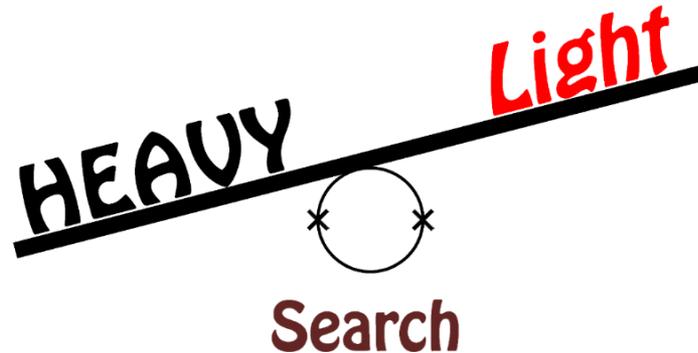


- Phase 1: 1.5 L at 4K - dip probe. Now!
- Phase 2: 30 L at 4K
- Phase 3: 1000 L at 100 mK
  - Multiplexed extension
- 1 month per decade

# Phase 1 Experiment Now Funded!



- Phase 1 funded by Kavli Foundation through Kavli Institute for Particle Astrophysics and Cosmology
- Design now underway; construction to begin in September
- Proposed name:

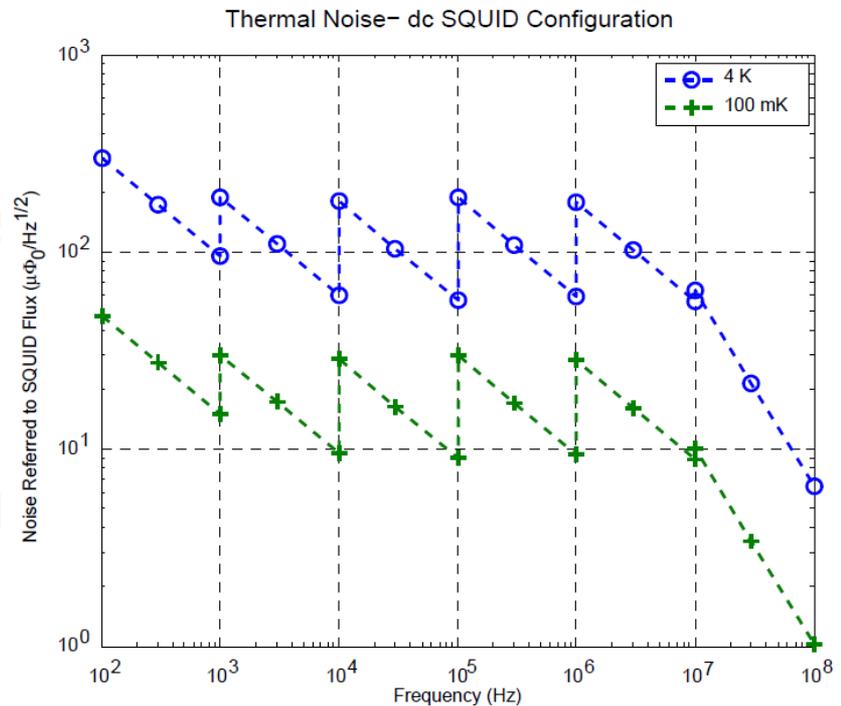


# Extra Slides



# $f < 10$ MHz: dc SQUID

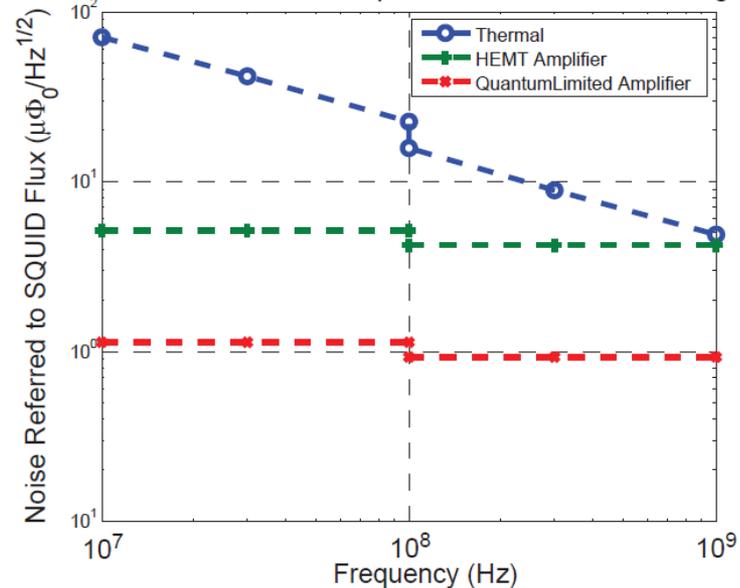
- Amplifier backaction from resistive shunts
- Upper limit set by noise degradation, also potentially parasitic capacitance between SQUID washer and input coil



# 10 MHz <math>f</math> <math>< 1</math> GHz: Microwave SQUID

- Dissipationless rf SQUID coupled to a microwave resonator
- rf SQUID acts as flux-dependent inductor- read out flux change as change in resonance frequency
- Sub-gap loss in junction instead of resistive shunts
- Largest readout noise contribution from microwave amplifier, e.g. HEMT or quantum-limited paramp
- >1 GHz and operating temperature ~100 mK: Quantum noise > thermal noise

Thermal Noise at 100mK and Amplifier Noise ac SQUID Configuration



# Acknowledgments

- Kavli Foundation and Kavli Institute for Particle Astrophysics and Cosmology (KIPAC) for Phase 1 funding
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